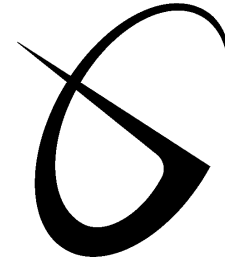


Global VSAT Forum



PERFORMANCE AND TEST GUIDELINES FOR TYPE APPROVAL OF 'COMMS ON THE MOVE' MOBILE SATELLITE COMMUNICATIONS TERMINALS

GVF-105

Released document.

This document defines the applicable performance requirements and test procedures for GVF Type Approval of "Comms on the Move" mobile satellite communications antenna systems.

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1 Purpose

The GVF type approval process is defined in detail in document GVF-101.

This document adds guidance for testing parameters that are unique to COTM terminals. These include the terminal tracking accuracy and the up-link inhibit functions should the tracking accuracy exceed published specifications. This testing is in addition to the test requirements outlined in GVF-101 [Mutual Recognition of Performance Measurement Guidelines and Procedures for Satellite System Operator Type approvals].

2 Terminal Definition

A COTM terminal is defined for the purposes of this document as follows:

- 1) C, X, Ku or Ka-Band operation.
- 2) Intended for operation on geostationary (non-inclined) satellites.
- 3) Automatically deploys, and accurately points, its antenna towards a designated target satellite.
- 4) Fully stabilized and suitable for vessels and moving vehicles.
- 5) Either (1) operates as part of a managed network, in which there is a hub or other counterpart earth station that is required for the terminal to operate and which supports terminal operations, or (2) operates as a standalone uplink station and all control functions are self-contained.
- 6) The type approval is granted to an agreed unique configuration, comprising the antenna (identified by manufacturer, model and diameter), mount, antenna control system, RF chain and modem. The maximum allowed EIRP density will be defined on the basis of the antenna RF performance and the results of the auto-deploy tests. The existing communication between modem and the platform shall be described by the manufacturer e.g. presence of power control in the modem, modem locking signal characteristics etc. In case of a different type of modem replacing the one of the originally type approved configuration, additional test plans will have to be mutually agreed between the applicant and the ATE (Authorized test Entity). If the manufacturer demonstrates that the pointing and RF parameters are not subject to a high level of modem dependency, other types of modems can be validated by analysis only.

3 Background

COTM applications place severe restrictions on the size of the antenna used as profile drag and limited footprint are key design elements for vehicles where these terminals are mounted. In many cases, this entails the use of Ultra Small Aperture Terminals (USATs) (i.e. < 1.0m at Ku-Band) that are inherently restricted in their focusing ability and thus produce wider beams. The combination of these wider beams with the increase risk of miss-pointing, due to the dynamic nature of the platform these terminals are mounted on, makes it critical to ensure that these terminals don't cause excessive levels of interference.

ETSI EN 302 977 and FCC 25.226 address the issue of adjacent satellite interference by imposing limits on the levels of off-axis emissions produced by the COTM terminal as well as requiring the terminal to inhibit transmission when a pointing error exceeding the manufacturer's declared threshold occurs as this would imply that the off-axis emissions limits are possibly being exceeded by the terminal if transmission is uninhibited. ETSI EN 302 977 also addresses the issue of on satellite cross-polarization interference that linearly polarized COTM terminals may cause due to tracking error on the antenna's transmit polarization axis.

These however don't provide test methodologies to verify compliance. It is therefore the intent of this document to present test methods, for conditions unique to COTM terminals, that would be required in addition to the general VSAT test requirements outlined in GVF-101.

4 Reference Performance Requirements

The summary specifications listed below are for reference only. Some satellite operators may require additional/different performance specifications.

4.1 ETSI EN 302 977 (Ku-Band COTM)

Off-Axis EIRP Spectral Density Limits:

- Co-Pol:
 - 33 - 25 log ($\varphi + \delta\varphi$) - H dBW/40KHz for $2,5^\circ \leq \varphi + \delta\varphi \leq 7,0^\circ$
 - 12 - H dBW/40KHz for $7,0^\circ < \varphi + \delta\varphi \leq 9,2^\circ$
 - 36 - 25 log ($\varphi + \delta\varphi$) - H dBW/40KHz for $9,2^\circ < \varphi + \delta\varphi \leq 48^\circ$
 - - 6 - H dBW/40KHz for $\varphi + \delta\varphi > 48^\circ$

- Cross-Pol:
 - 23 - 25 log ($\varphi + \delta\varphi$) - H dBW/40KHz for $2,5^\circ \leq \varphi + \delta\varphi \leq 7,0^\circ$
 - 2 - H dBW/40KHz for $7,0^\circ < \varphi + \delta\varphi \leq 9,2^\circ$

Notes:

- φ is the angle, in degrees, between the main beam axis and the direction considered
- $\delta\varphi$ is the pointing error threshold, in degrees, as declared by the applicant
- H (in dB) is $10\log N$ where N is the maximum number of VMESs which may transmit in the same carrier frequency band as declared by the applicant.
- The radiation pattern planes for which these restrictions apply can be limited if documentary evidence is provided that not all directions are concerned. TR 102 375 gives guidance on determination of the concerned subset.

Pointing Error and Cessation of Emissions:

- The maximum pointing error requirement is to be set by the applicant as long as the terminal meets their off-axis EIRP spectral density limits when mispointed by this value.

- The terminal shall cease transmission if the maximum pointing error is exceeded for a period of T seconds (T is set by the applicant).
- The terminal will not restart transmission unless the pointing error is within the limit for a period of 2xT seconds.
- T is not to exceed 2 seconds

4.2 FCC 25.226 (Ku-Band COTM)

Off-Axis EIRP Spectral Density Limits:

- Co-Pol in the GSO plane:
 - $15-10\log(N)-25\log\Theta$ dBW/4kHz for $1.5^\circ \leq \Theta \leq 7^\circ$
 - $-6-10\log(N)$ dBW/4kHz for $7^\circ < \Theta \leq 9.2^\circ$
 - $18-10\log(N)-25\log\Theta$ dBW/4kHz for $9.2^\circ < \Theta \leq 48^\circ$
 - $-24-10\log(N)$ dBW/4kHz for $48^\circ < \Theta \leq 85^\circ$
 - $-14-10\log(N)$ dBW/4kHz for $85^\circ < \Theta \leq 180^\circ$

For $\Theta > 7^\circ$ the envelope can be exceeded by up to 10% of the sidelobes by a maximum of 3dB

- Co-Pol in all other planes:
 - $18-10\log(N)-25\log\Theta$ dBW/4kHz for $3.0^\circ \leq \Theta \leq 48^\circ$
 - $-24-10\log(N)$ dBW/4kHz for $48^\circ < \Theta \leq 85^\circ$
 - $-14-10\log(N)$ dBW/4kHz for $85^\circ < \Theta \leq 180^\circ$

Envelope can be exceeded by up to 10% of the sidelobes by a maximum of 6dB

- Cross Pol:
 - $5-10\log(N)-25\log\Theta$ dBW/4kHz for $1.8^\circ \leq \Theta \leq 7.0^\circ$
 - $-16-10\log(N)$ dBW/4kHz for $7.0^\circ < \Theta \leq 9.2^\circ$

Pointing Error and Cessation of Emissions:

- Requires compliance to one of the following 2 conditions regarding maximum pointing error, each of which has a specific cessation of emission requirement:
 - A. Maximum pointing error $\leq 0.2^\circ$
 - Cessation of emission occurs within 100 msec if pointing error exceeds 0.5°
 - Transmission to resume when pointing error $\leq 0.2^\circ$

Or

- B. Determined by the applicant based on the terminal still meeting the FCC off-axis EIRP density limits when mispointed by this value (Similar to the ETSI definition)

- Cessation of emissions occurs within 100 msec if pointing error exceeds specified pointing error threshold
- Transmission to resume when pointing error is below threshold

4.3 On-Satellite Cross-Pol EIRPSD Limit References

Within the stated pointing error threshold of the terminal, the EIRPSD of the cross-polarized signal should not exceed:

- For Ku-Band Terminals:
 - 0 dBW/4KHz (Equivalent to a 1.2m antenna compliant with FCC 25.209 and FCC 25.218 for digital Ku-Band earth-station operation, having an on satellite Cross-Pol isolation of 30dB).
 - 8 dBW/4KHz (Equivalent to a 1.2m antenna compliant with FCC 25.209 and ITU-R S.728-1, having an on satellite Cross-Pol isolation of 30dB).

4.4 Antenna Radiation Pattern Limit Reference

In addition to the limit on EIRPSD, some satellite operators may require a limit on the antenna gain radiation pattern. One such limit can be found in ITU-R S.465-6 summarized below:

- $\text{Gain} \leq 32 - 25 \log \Theta$ dBi for $\Theta_{\min} \leq \Theta \leq 7^\circ$

Where:

$\Theta_{\min} = 1^\circ$ or $100 \lambda/D$ degrees, whichever is the greater, for $D/\lambda \geq 50$.

$\Theta_{\min} = 2^\circ$ or $114 (D/\lambda)^{-1.09}$ degrees, whichever is the greater, for $D/\lambda < 50$.

It should be noted that for non-rotationally symmetric apertures, D is the dimension of the aperture along the plane of the antenna radiation pattern being assessed.

5 Type Approval Test Requirements

Type approval for COTM terminals shall include the following provisions:

- a) COTM Terminals employing Traditional VSAT antennas: Antennas included in this classification shall be tested per the requirements of this document [GVF-105] and GVF-101 in general and section 4 of GVF-101 in particular.
- b) COTM Terminals employing ESA technology or other solutions in which the antenna performance (pattern, gain, polarization and G/T) changes as the antenna pattern is scanned off axis: COTM terminals falling in this classification require special test procedures in addition to the requirements of this document [GVF-105]. This is because of the volume of test measurements that would otherwise have to be undertaken if standard measurement procedures were employed. In summary, the pattern characteristics of any COTM terminal falling in to this grouping shall include the following elements. These procedures are outlined in section 4.4 of GVF-101. *[GVF-101 is being updated to include Sect 4.4 as an addition.]*

- 1) Provide predicted, principal plane, co- and cross pol antenna pattern coverage plots at the lower, mid and upper frequencies in the transmit and receive bands of interest.
- 2) Conduct physical antenna pattern measurements corresponding to the pattern predictions for the step (1) above.
- 3) Compare predicted vs. measured pattern coverage to validate the accuracy of the pattern prediction software.
- 4) Use the pattern prediction software to identify other regions where specification violations may occur or for areas with minimal performance margins.
- 5) Conduct physical pattern measurements over the same angular and skew angle regions identified in step (4) above.

The rationale for this approach is simple. First, the accuracy of the prediction software requires validation by comparing predicted performance against measured performance. Second, having established the accuracy of the predicted performance, the process moves to characterize the actual antenna performance for conditions where the predictions indicate areas of risk or non-compliance.

- c) COTM Terminals employing Radomes: Radomes can degrade the performance of the antenna being protected by the radome. The provisions of GVF-105 and GVF-101 shall require that any terminal [COTM, Auto-Deploy or Stationary terminal] normally supplied with a radome as part of its Bill of Materials, be characterized with the radome in place. The radome shall be positioned with respect to the terminal antenna as would be encountered in normal terminal operation. This provision shall apply to all antennas falling on the classifications of paragraphs a) and b) above.

5.1 Pre-defined parameters

The manufacturer or type approval applicant must propose the following parameters:

- 1) Maximum BPE under normal operating conditions†.
- 2) The length of time T that when the BPE defined above is exceeded for will cause a cessation of emission.
- 3) For linearly polarized terminals, the cross-pol isolation within the BPE of the terminal.
- 4) Measured antenna radiation patterns for both Co and Cross polarized components showing an angular range of +/- 7° from bore-sight with a resolution better than 0.1°:
 - a. For circular aperture antennas these should include 2 orthogonal cuts (e.g. AZ and EL) for 2 orthogonal polarization settings (e.g. Vertical and Horizontal)
 - b. For elliptical apertures these should include 19 cuts, starting with the AZ cut (i.e. 0° skew cut) going thru to the EL cut (i.e. 90° skew cut) in 5° increments. It is important to note that the maximum compliant on-axis EIRPSD for COTM terminals with non-aligning elliptical aperture antennas will depend on the skew angle of the operating scenario (See appendix for example).

5.2 Required Tests

In addition to testing defined in GVF-101 (VSAT level), COTM terminal testing should include:

- The verification that the manufacturer's stated pointing error threshold is not exceeded under normal operating conditions†.
- The verification of the "transmit inhibit" function when the stated pointing error threshold is exceeded for the stated time T.

- The verification of the level of Cross-Pol interference that the terminal emits towards the desired satellite under normal operating conditions†.
- The verification of antenna pattern "skew" corrections for non-circular apertures.

†Normal operating conditions are minimum profiles to be agreed upon between the terminal manufacturer and the approved GVF test entity based on the target customer application for the terminal. A description of these profiles is to be included in the type approval certificate.

These tests, in combination with the terminal's radiation pattern measurement (per GVF-101) will allow for determination of a maximum compliant level of on-axis EIRP density for the terminal as shown in Figure 1.

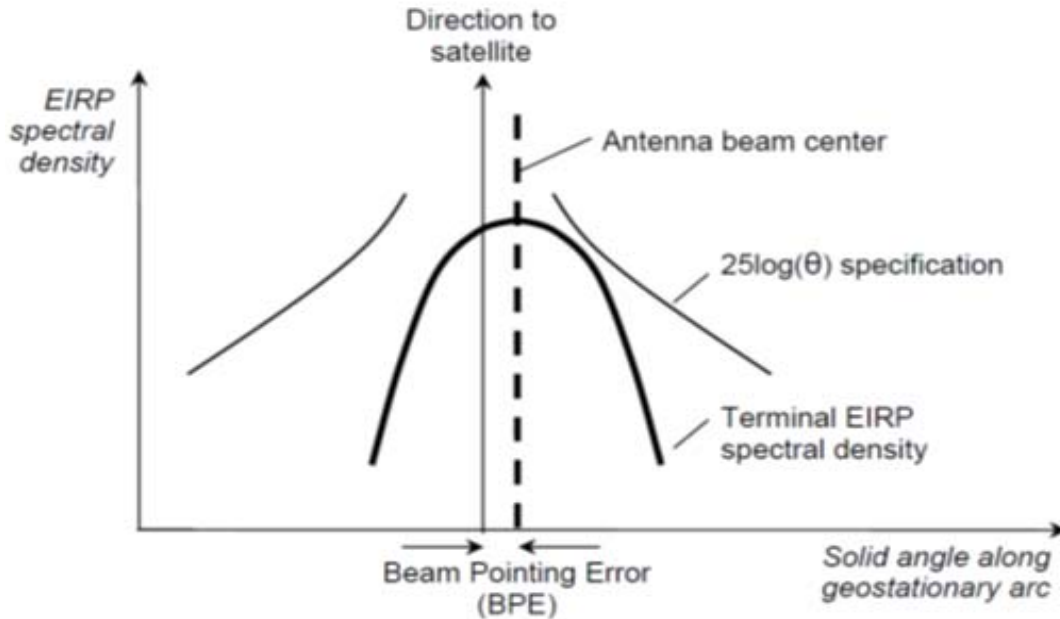


Figure 1: Determination of maximum compliant on-axis EIRP density

5.3 Test Methods

5.3.1 Pointing Accuracy Testing of Terminals with Small Antennas (ie. $\leq \sim 0.8\text{m}$ Diameter at Ku-Band)

Typically, for terminals with large antennas where gain of the main lobe quickly varies away from boresight, pointing accuracy can be accurately determined by monitoring the signal level received by the target satellite since fluctuation of this signal is dominated by changes in pointing error.

When dealing with terminals with small apertures, as is the case for many COTM terminals, the signal fluctuation at the target satellite caused by pointing error becomes very small, due to the beam flatness near the beam peak, making accurate measurement of pointing error impractical using this approach.

The alternative herein described utilizes the fact that these small antennas have a radiation pattern main lobe that is wide enough to be detected at an adjacent satellite providing signal levels suited for calibrated measurements.

Essentially, a signal is uplinked to the primary satellite of interest and the downlink signal is simultaneously monitored on both the primary and adjacent satellites. The COTM VSAT terminal is then driven over a course, suggested by the manufacturer and agreed upon by the GVF test entity, which will be representative of typical use of the terminal. The COTM VSAT terminal can equally be installed on a motion table set up to replicate movement of the vehicle. When monitoring the downlink signal on the primary satellite, very small (and often hard-to-detect) changes in received strength may be seen due to small pointing errors as the terminal antenna attempts to compensate for vehicle motion. However, the changes in signal strength recorded on the adjacent satellite downlink will be amplified due to the rapidly changing gain for small angular errors. Once differences between propagation paths (SFD, Receiving aperture G/T ...) have been corrected for, the difference between both measured signals referred to as Δ will correspond to the antenna radiation pattern. Since the adjacent satellite is around 2 degrees offset from the beam peak this Δ is quite sensitive to pointing error.

Δ (dB) = Measured received signal from adjacent satellite - Measured received signal from desired satellite + Correction Factor

Given linear operation of the Terminal under Test, the link via the primary satellite of interest, and the adjacent satellite, it can be shown that the following relationships exist:

$$P_{rcv\ Sat1} = P_{up\ Sat1} - L_{up\ Sat1} + G_{Sat1} - L_{downSat1} + G_{Mon\ ET1}$$

and

$$P_{rcv\ Sat2} = P_{up\ Sat2} - L_{up\ Sat2} + G_{Sat2} - L_{downSat2} + G_{Mon\ ET2}$$

Where:

- $P_{rcv\ Satx}$ = Power received from Satellite x , either primary(1) or adjacent(2)
- $P_{up\ Satx}$ = Power uplinked from Terminal Under Test towards Satellite x
- $L_{up\ Satx}$ = Uplink loss from Terminal Under Test towards Satellite x
- G_{Satx} = Gain through Satellite x transponder
- $L_{down\ Satx}$ = Downlink loss from Terminal Under Test towards Satellite x
- $G_{Mon\ ETx}$ = Gain through Measurement Earth Terminal x

With linear operation of links via the primary and adjacent satellites during the period of measurement, the Correction Factor can combine the individual Loss and Gain terms into one factor. If the Correction Factor is time varying, it might still be possible to combine into one term, but with added measurement complexity. Correction Factor can be calculated by determining that value for which:

$$F_{correction} = P_{rcv\ Sat1} - P_{rcv\ Sat2} - G_{up\ TUT\ at\ spacing}$$

Where:

- $F_{correction}$ = Correction Factor
- $G_{up\ TUT\ at\ spacing}$ = Gain difference in uplink antenna pattern of Terminal Under Test at the offset angle of the orbital spacing of Satellite 1 and Satellite 2.

An example of a radiation pattern of a terminal under test along the GSO arc is shown in Figure 2. It can be seen that when the pointing error is 0° , a $\Delta = -16.7\text{dB}$ should be measured. When a pointing error of 0.5° towards the adjacent satellite occurs, an increase to $\Delta = -8.2\text{dB}$ will be recorded.

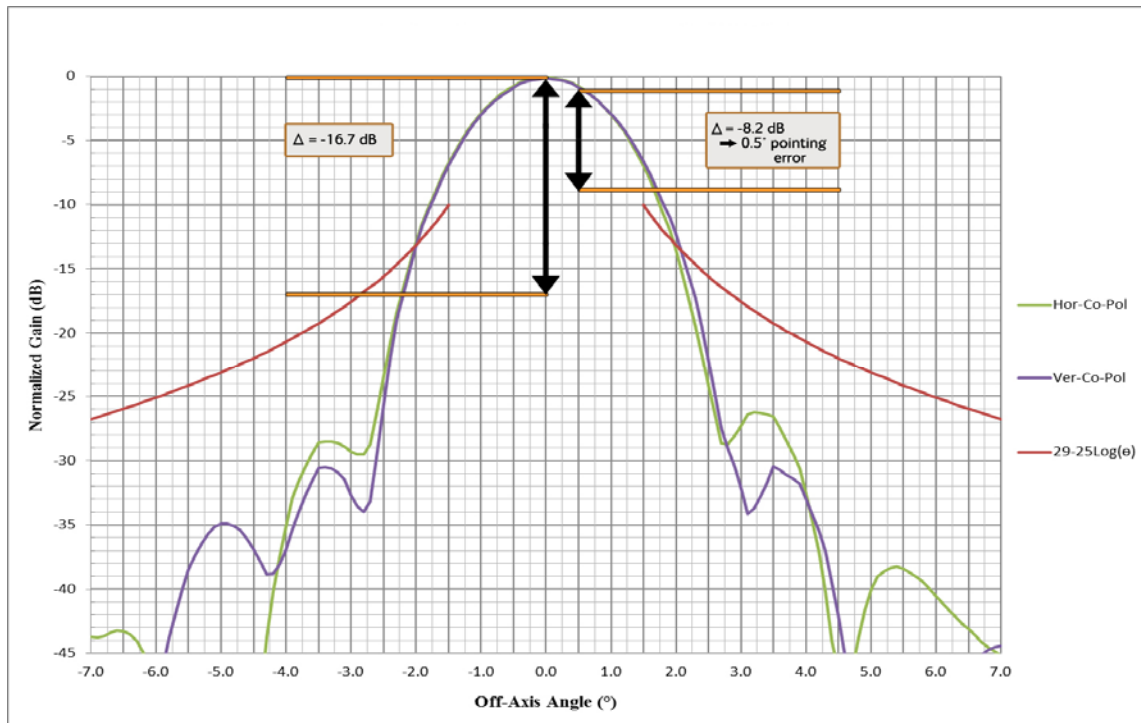


Figure 2: Typical Radiation pattern for COTM terminal with small antenna

The correction factor can be obtained while the terminal is static, by manually peaking the antenna using the -6dB points. The difference between the expected Δ (obtained from the radiation pattern and knowledge of the adjacent satellites relative position with respect to bore-sight) and the measured one while the terminal under test is peaked corresponds to the correction factor.

The essential element of this Test Plan is to operate a mobile VSAT terminal under controlled conditions such that:

- Mobile VSAT transmit antenna input power spectral density is maintained as constant as possible.
- Calibrated measurements are made simultaneously on the downlink power spectral density of the desired satellite and the downlink power spectral density of the adjacent satellite or satellites.

Assuming that the entire satellite communications link remains within the linear region on both the satellite of interest and the adjacent satellites, any impact on EIRP spectral density will result in a direct dB for dB downlink spectral density change. Simultaneous measurements of the downlink spectral density from each of the satellites should result in data which can then be converted from EIRP spectral density to antenna angular offsets. It is that principle which forms the basis of the Test Plan presented here.

The details of the test procedure may be adapted to accommodate a particular terminal configuration, however "live" on-air test presents an effective method to accurately determine the pointing accuracy for a VSAT terminal when in motion.

Using this approach, the pointing accuracy stated by the manufacturer can be verified thru several trials where the terminal is operated under conditions mimicking standard user operation. A typical processed data capture from such a trial is shown in Figure 3.

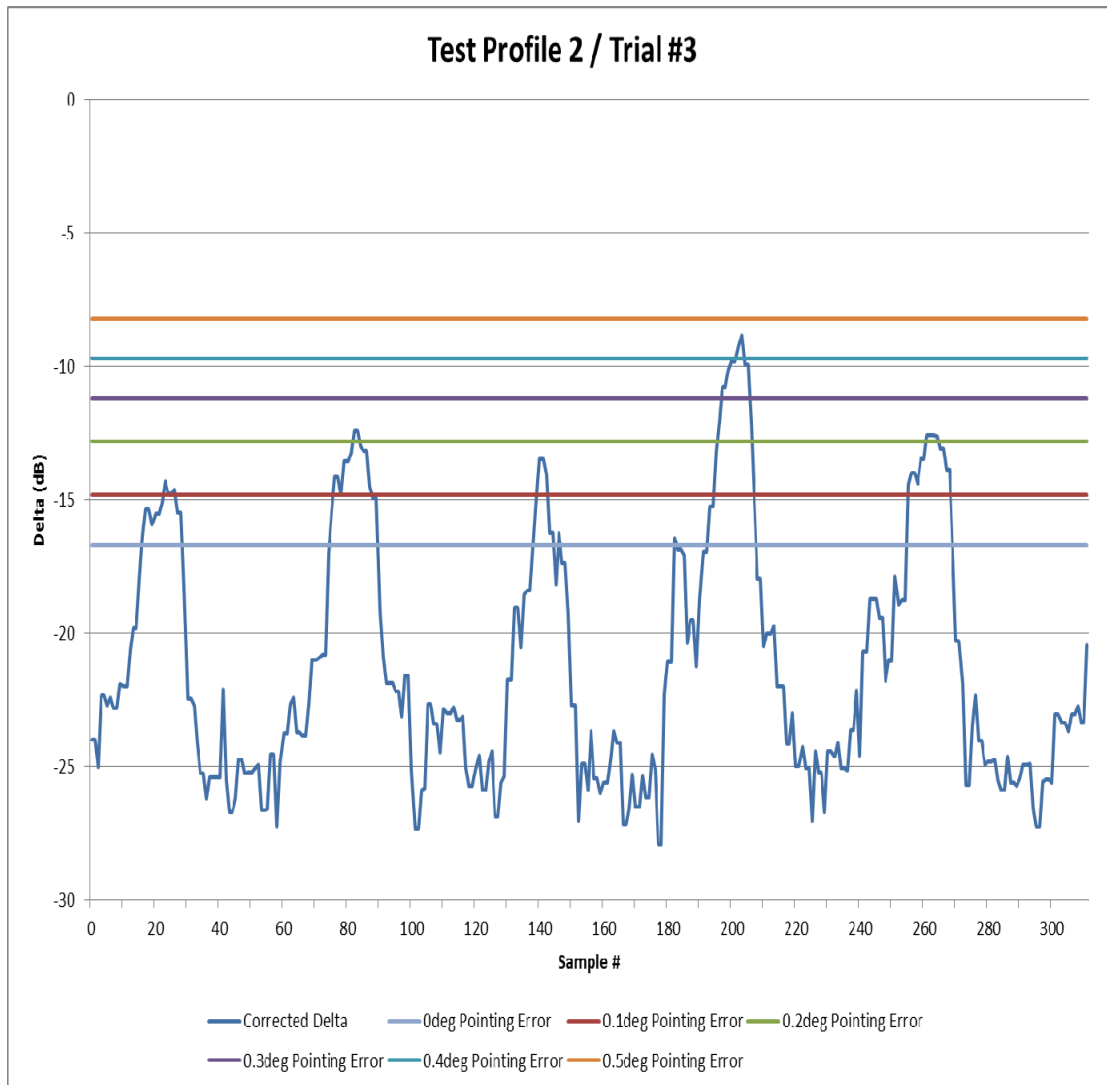


Figure 3: Typical processed data plot of a trial of the pointing accuracy verification

5.3.2 Verification testing of the “Transmit Inhibit” function

Using the same method of measuring pointing accuracy described in 5.3.1, it is possible to verify the transmit inhibit parameters (pointing error threshold and time T) stated by the manufacturer by repeatedly inducing a pointing error that exceeds the error threshold for a period longer than T. Figure 4 shows a typical processed data capture of this testing. In this case, the system inhibits transmit when a pointing error threshold of 0.5° is exceeded for $>100\text{ms}$. As can be seen, when the system is muted, Δ goes to $\sim 0\text{dB}$ (this is not exactly 0 because of the noisiness of the measured signal when the terminal is muted).

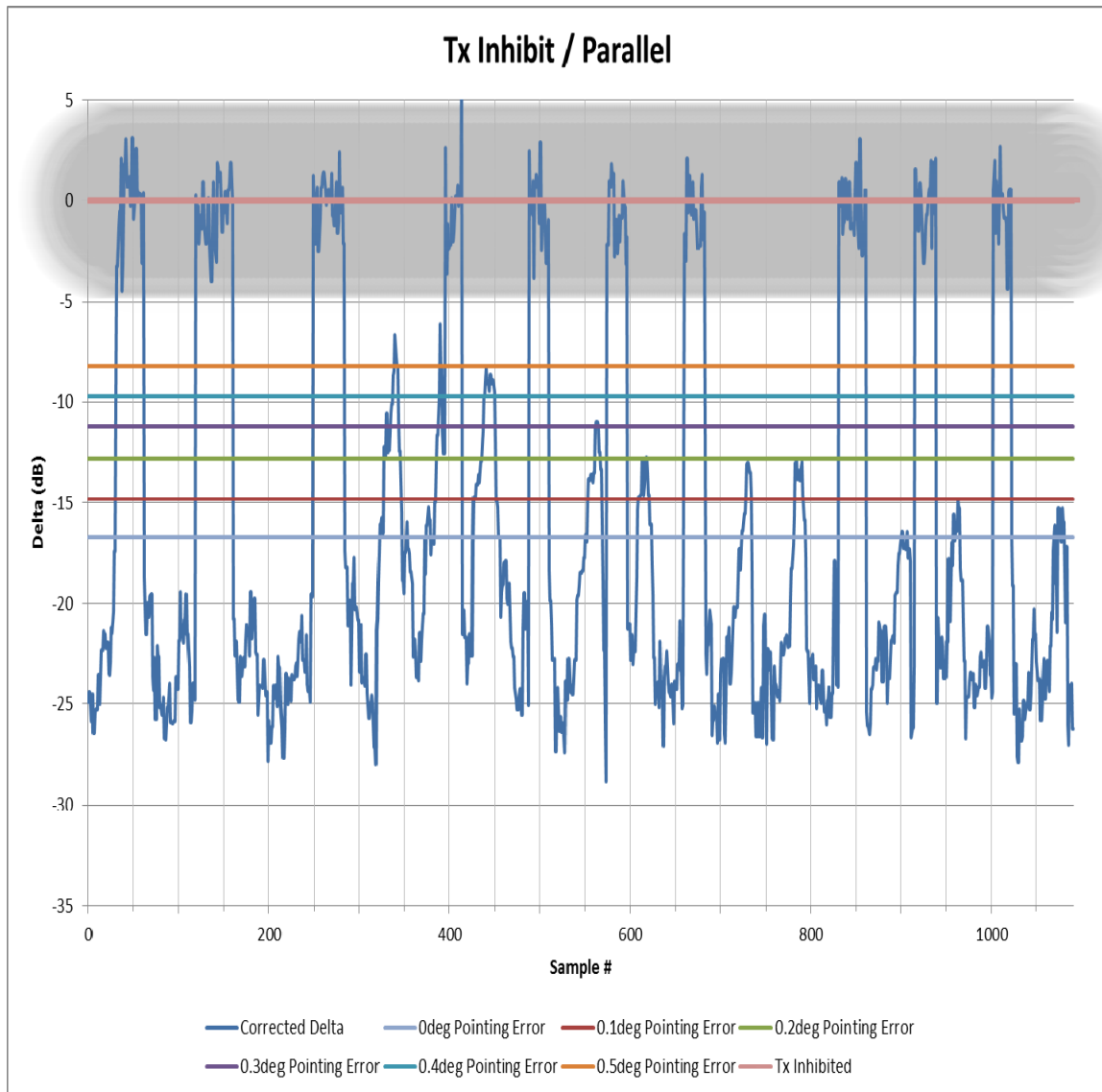


Figure 4: Typical processed data plot of a profile used for transmit inhibit verification

5.3.3 Verification testing of the level of Cross-Pol interference that the terminal emits towards the desired satellite under normal operating conditions

The purpose of this test is to verify the level of Cross-Pol interference emitted by the terminal towards the desired satellite under normal operating conditions relative to the peak Co-Pol signal level that can be seen on the desired satellite for the same terminal settings and EIRP level.

The test should be conducted in under clear sky conditions. The test starts by manually (i.e. using a “jog” type function on the controller) peaking the terminal by finding the -6dB points. Next transmission from the terminal under test is established. The Co-Pol signal received by the desired satellite is monitored by the hub station for a few minutes. The maximum received signal level is recorded. The COTM terminal is then operated under the agreed upon operating profiles while the hub station is monitoring the Cross-Pol level recorded. Figure 5 shows a typical processed data capture of this testing.

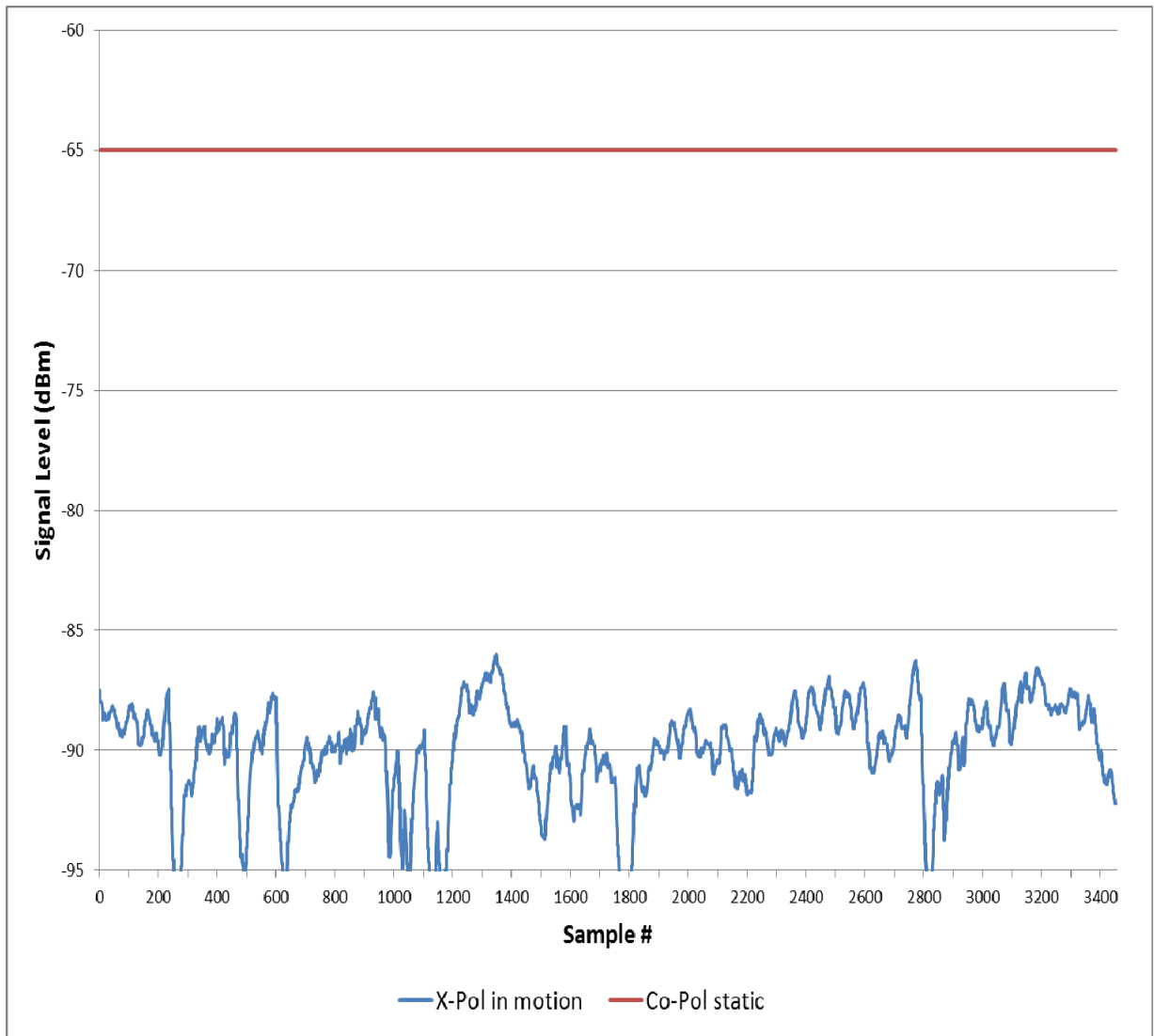


Figure 5: Typical processed data plot of a profile used for Cross-Pol Interference level testing.

5.3.4 Verification testing of antenna "skew" corrections for non-circular apertures

COTM terminals can utilize antenna apertures having circular or non-circular radiation patterns. Circular apertures exhibit radiation patterns which are essentially equal in all directions away from the center of the beam. Non-circular apertures, elliptical or other shapes, exhibit narrower effective beamwidths along the longest axis of the aperture. They exhibit wider effective beamwidths along the narrowest axis. This results in a "skew" of non-circular radiation patterns along the geostationary arc which must be considered as it relates to ASI. An example of the definition of antenna radiation pattern "skew", using a rectangular aperture radiation pattern, is provided in Figure 6.

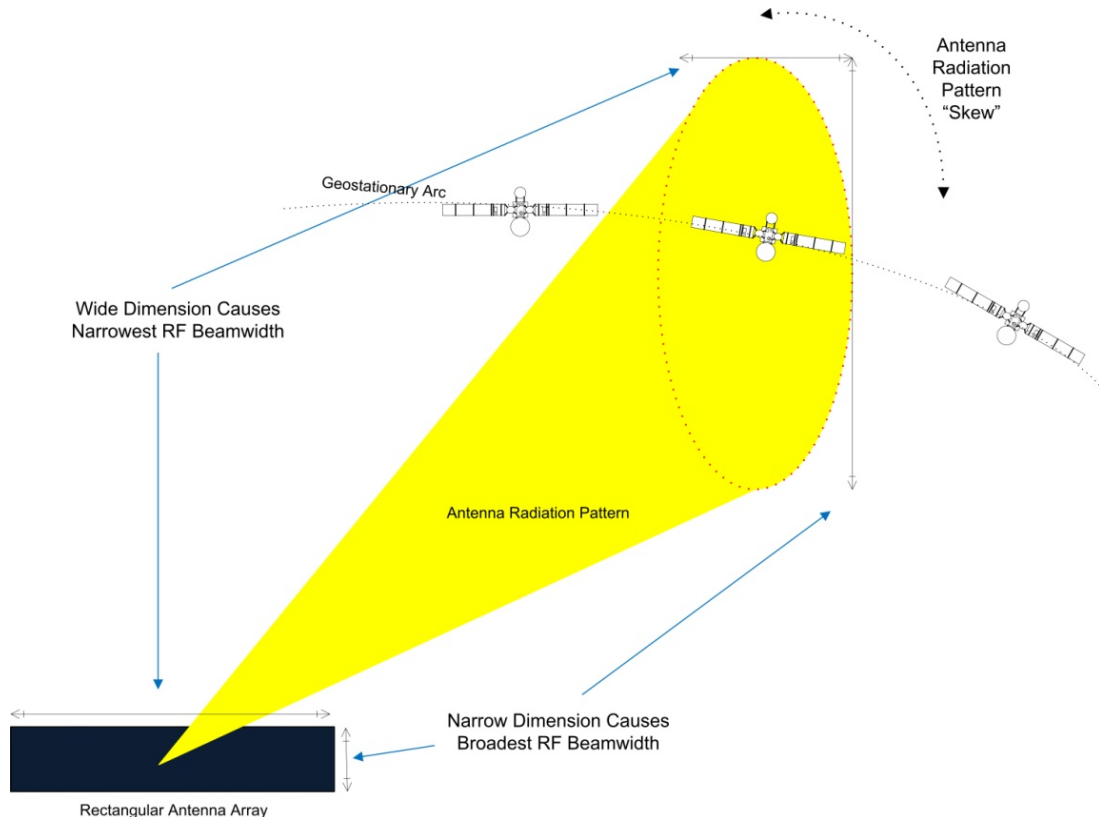


Figure 6: Non-circular antenna aperture radiation pattern "Skew".

COTM terminal antenna pattern "skew" will be present in terminals utilizing non-circular apertures. The effects of antenna radiation pattern Skew can be compensated with one of three methods:

- 1) Establish the COTM terminal's EIRP power spectral density at a level which will satisfy the ASI requirements and radiation in directions other than the geostationary arc under the worst condition.
 - Terminals employing this method require no skew angle compensation testing.
- 2) Establish the COTM terminal's EIRP power spectral density at a level which is intended to satisfy ASI requirements and radiation in directions other than the geostationary arc up to a defined maximum skew angle, and for higher Skew angles to mute the COTM transmitter.
 - Terminals employing this method require verification testing of this mute functionality. This can be performed using the following procedure:

- a. Mount a 2 axis inclinometer on the azimuth rotating part of the antenna while aligning one axis such as it is parallel to the elevation axis (Pitch) and the other parallel to the beam direction if the antenna were pointed at 0° elevation (roll).
- b. Mount the terminal to a base.
- c. Acquire a satellite that is within the terminal's allowable skew range but close to the manufacturer's maximum stated value.
- d. While monitoring the RF input power of the antenna (using BUC M&C function or a coupler...), gradually tilt the base to which the terminal is mounted until the antenna RF input power is seen to drop to the noise floor of the test setup.
- e. Note the Pitch and Roll values at which the transmission was muted.
- f. Using the following equations from [1] (where Pol corresponds to skew) calculate the corresponding skew of the tilted antenna at which transmission was muted and compare with the manufacturer's stated value:

$$Pol_{Tilted} = Pol_{Untilted} - a \cosd \left(\frac{\sqrt{(\cosd(\tau))^2 - (\sind(Elevation_{Untilted}))^2}}{\cosd(Elevation_{Untilted})} \right)$$

Elevation Correction for tilted applications

Where:

$$\tau = -a \operatorname{atand}(\cosd(Pitch_{ant}) \operatorname{tand}(Roll_{ant}))$$

- 3) Dynamically adjust the COTM terminal's EIRP power spectral density using a determination of the radiation patterns Skew along the geostationary arc, where it affects ASI, and in directions other than the geostationary arc as defined in regulatory requirements.
 - Terminals employing this method require verification testing of this functionality. This can be performed using the following procedure:
 - a. Mount a 2 axis inclinometer on the azimuth rotating part of the antenna while aligning one axis such as it is parallel to the elevation axis (Pitch) and the other parallel to the beam direction if the antenna were pointed at 0° elevation (roll).
 - b. Mount the terminal to a level base.
 - c. Acquire a satellite and set the antenna's input PSD to within 1 dB of the allowable limit for the corresponding location/satellite combination.
 - d. Tilt the base so as to increase the terminal skew.
 - e. Note the input PSD to the antenna at the new terminal skew.
 - f. Using the above equations from [1] (where Pol corresponds to skew) calculate the corresponding skew of the tilted antenna.
 - g. Compare the on-axis EIRP density of the tilted terminal to the corresponding Skew specific EIRP density value specified by the manufacturer.

Other potential verification test methodologies could be utilized to verify the radiation pattern skew compensation process, subject to minimum profiles to be agreed upon between the terminal manufacturer and the approved GVF test entity based on the target customer application for the terminal.

Using any methodology, results of verification tests and any imposed operational constraints of antenna skew compensation should be documented for COTM terminals utilizing non-circular antenna apertures

6 References

- [1] Technische Universität München – Institute of Astronautics in Co-operation with Eutelsat Communications “Calculation of Azimuth, Elevation and polarization for non-horizontal aligned Antennas, Analytic Formulas”, Technical Document TD-1205-b.

Appendix A: Additional Notes on Pointing Accuracy Verification Method

EIRP spectral-density levels anticipated from various different mobile VSAT antenna apertures and the FCC EIRP spectral-density limits are graphically illustrated in Figure 7 below. In this case, the individual mobile VSAT antenna is presumed to be perfectly pointed towards the desired satellite, thus shown as the 0 Degrees angle in Figure 7. Input power spectral-density for each of the different mobile VSAT apertures illustrated in Figure 7 were adjusted to a maximum value of -22 dBW/ 4kHz. The FCC VMES EIRP spectral-density limit is similarly shown graphically for comparison.

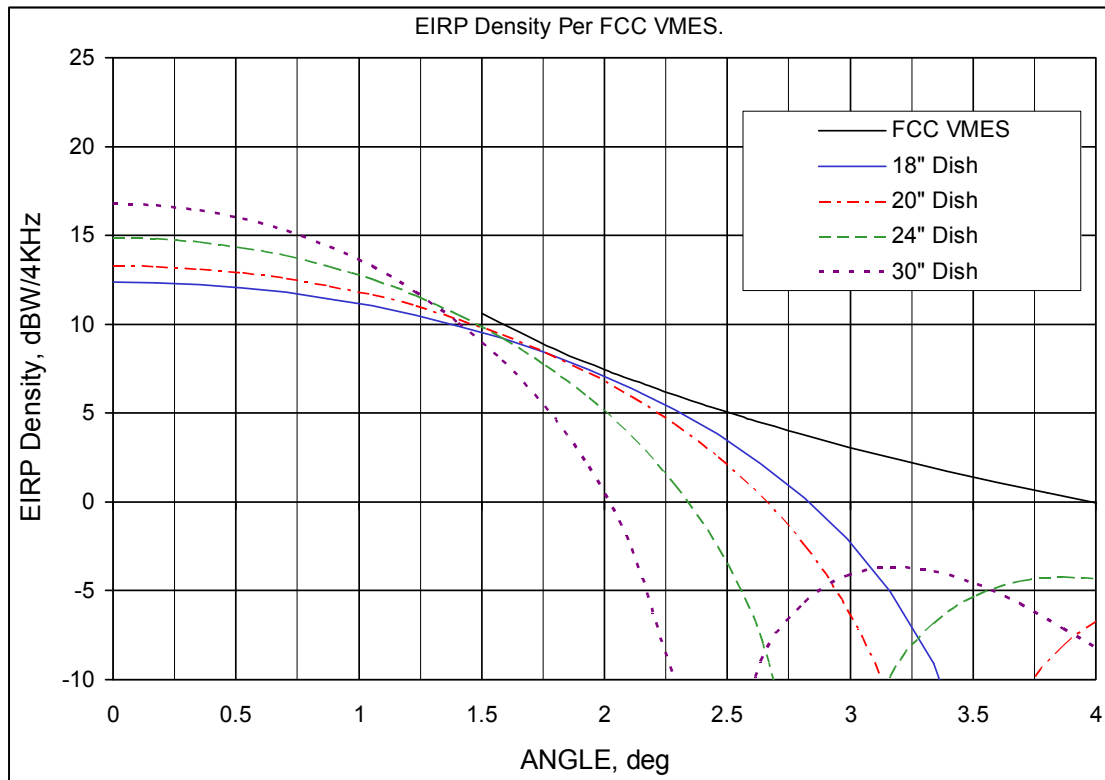


Figure 7 VMES /mobile VSAT EIRP spectral-density comparison

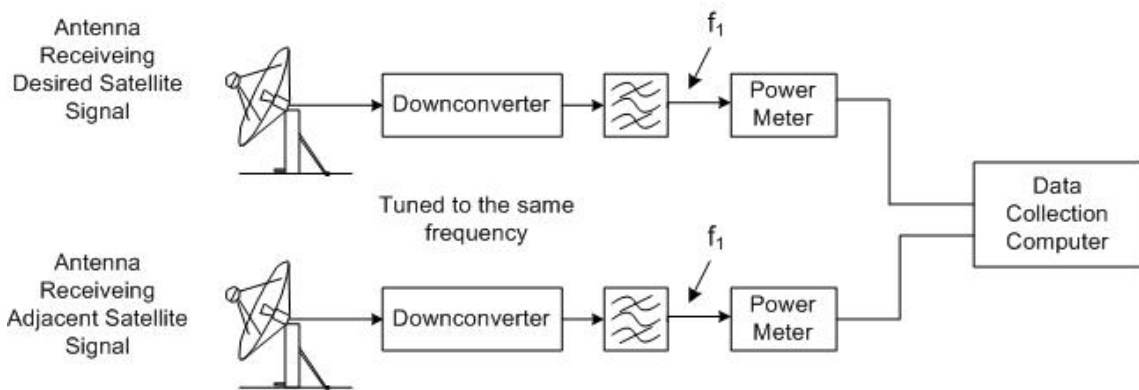
Using the 24-inch aperture as an example, we can describe the expected measurement performance during the test. On the satellite of interest, this antenna should exhibit a maximum EIRP spectral-density of approximately 15 dBW/4 kHz. Simultaneously, an adjacent satellite spaced 2 Degrees away on the GSO arc should see an EIRP spectral-density of approximately 5 dBW/4 kHz. If the mobile VSAT antenna were to remain perfectly pointed these levels should remain constant, within the measurement tolerances driven by the accuracies of the test system and any propagation fade conditions. If the antenna was to be pointed towards the adjacent satellite by 0.25 Degrees, the on-axis EIRP spectral-density would degrade by approximately 0.1 dB or less, but the EIRP spectral-density towards the adjacent satellite would increase by approximately 3 dB. Similarly, if the antenna was to be pointed away from the adjacent satellite by 0.25 Degrees, the on-axis EIRP spectral-density would degrade by approximately the same 0.1 dB or less, but the EIRP spectral-density towards the adjacent satellite would decrease by approximately 3.5 dB.

From Figure 7 it can be seen that this same approach could be utilized to measure the difference in EIRP spectral-density, and thus calculate angular offset, for any of the mobile VSAT aperture sizes under consideration. Larger antennas tend to have higher EIRP spectral-density slopes at

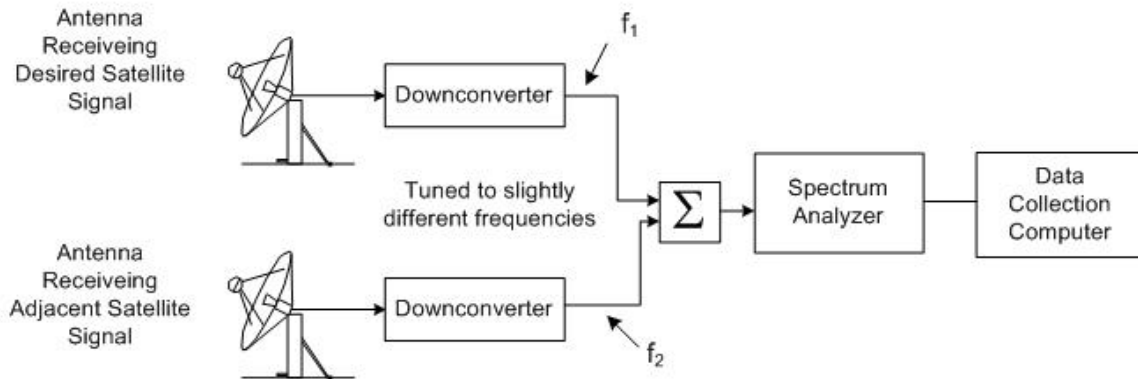
2 Degree offsets and smaller antennas have lower EIRP spectral-density slopes. The accuracy in the ability to calculate angular offsets for a particular aperture is thus a function of the accuracy and sensitivity in EIRP spectral-density measurements driven by the test conditions and the measurement equipment used.

By making careful simultaneous measurements of both the EIRP spectral-density on the satellite of interest and the EIRP spectral-density on one adjacent satellite sufficient resolution should be available to determine differences on the order of 0.1 dB or less. These changes in EIRP spectral-density can then be converted to angular offset errors in degrees by comparing the values measured with the antenna patterns that would drive such values. Factors affecting the accuracy of the measurements and the calculated angular offsets will be described below.

There are at least two ways such measurements could be performed—using spectrum analyzer or power meter measurements. Figure 8 illustrates the basic block diagrams of the architecture that could be used in these two approaches.



A) Power Meter Measurement Approach



B) Spectrum Analyzer Measurement Approach

Figure 8 Antenna Pointing Measurement Architectures

To understand the relative merits of each of these architectures and make an appropriate selection for the tests to be conducted it is important to consider the potential measurement accuracy, factors affecting the measurement accuracy, and the relative cost and complexity of each.

A.1 Signal + Noise Measurements

The downlink signals present at each downlink antenna will consist of the desired signal plus noise, or (C+N). The objective of the measurement in each case is to measure the signal, or Carrier power spectral-density levels, C, and not the noise. Noise is present in each of the downlink signals so accurate measurements of Carrier power spectral-density can only be made if the noise energy is sufficiently low as to be insignificant to the Carrier measurements or it must be mathematically removed.

Power meters measure all the RF power, P, within their detector's broad frequency coverage range. In general, that amounts to several GHz of potential sensitivity bandwidth. Spectrum analyzers combine a filter and detector function in one package. They make several (C+N) measurements, always reduced to power spectral-density because they use a pre-detector filter to limit the power P arriving at the detector, with each sweep. Thus, with the Power Meter method we begin with Power measurements and with the spectrum analyzer we begin with power spectral-density measurement. However, with either technique we are faced with the need to reliably measure Carrier power spectral-density while eliminating the effects of Noise power spectral-density.

The downlink signals of interest can be considered to simultaneously consist of:

$$\text{Downlink Power} = P = C+N \quad (\text{Equation 1})$$

Where
P = Power (observed at the measurement device)
C = Carrier Power (the signal desired to be measured)
N = Noise Power (which must be eliminated)

Noise power present in the downlink signal can be further characterized as a specific Noise Power spectral-density in the bandwidth of interest. (In the case of a Power Meter, the bandwidth of interest is, unfortunately, the entire detector sensitivity bandwidth.) This relationship can be shown mathematically as:

$$\text{Noise Power} = N = N_0 \times B \quad (\text{Equation 2})$$

Where
N = Noise Power (which must be subtracted for accuracy)
N₀ = Noise Power spectral-density
B = Bandwidth (of the measurement bandwidth)

The value desired for measurement is the Carrier Power, C. Since Equation 1 demonstrates that the value observed is actually Downlink Power, C+N, we must mathematically remove the contribution of Noise Power, N, to make the desired measurement. This leads to our true measurement objective being:

$$\text{Carrier Power} = C = (C+N) - N \quad (\text{combining from Equation 2 produces})$$

$$C = (C+N) - N_0 \times B \quad (\text{Equation 3})$$

Where
C = Carrier Power (the signal desired to be measured)
(C+N) = Carrier plus Noise power (observed in measurement)
N₀ = Noise Power spectral-density
B = Bandwidth (of the measurement bandwidth)
N = Noise Power (which must be eliminated)

A.2 Consideration of the Power Meter approach

Power Meters are designed to make very accurate absolute RF power measurements. With the very broad bandwidth associated with Power Meter detectors it would be virtually impossible to make accurate measurements of the downlink signal using a Power Meter directly. The reason

for that is simply that even a very low value of N_0 would be multiplied by the very broad measurement bandwidth and it would then be difficult to utilize Equation 3 above to calculate the Carrier Power C . They are not designed to make power spectral-density measurements but can make such measurements with the aid of a pre-detector filter as shown in Figure 8 A above. As illustrated in Figure 8 A, a filter can be inserted ahead of the Power Meter detector for the specific purpose of reducing the bandwidth B affecting the $(C+N)$ measurement values. As bandwidth B is reduced to a very small value, and the Carrier Power C remains very narrow in frequency such as resulting from a CW signal, the contribution of C in Equation 1 above becomes much more significant than N . The complications induced in the measurements by utilizing such a filter are:

- measurement accuracy/calibration of each Power Meter detector must remain well-matched to ensure maximum accuracy.
- gain stability of the downconverter in each measurement leg can affect accuracy
- absolute bandwidth accuracy of each filter must be well-matched to ensure maximum accuracy
- for very narrow filter bandwidths, the center frequency of each signal must remain reliably within the filter passband to ensure maximum accuracy

For absolute accuracy, C is determined using Equation 1 and Equation 3 above. However, for the cases where C is significantly larger than N , or $C \gg N$, then the effect of N is lower than the measurement system accuracy. One can mathematically consider the measurement impact Noise contributions would have. With the Power Meter approach, it is very difficult to generate a strong enough $(C+N)$ signal as to make the contribution of N insignificant. A good test of such conditions is to eliminate the C energy, by turning off the transmit signal, and observing the resulting N component on its own. A value of N that is significantly below the value of $(C+N)$ can be shown mathematically to have an impact on the value of C because the bandwidth B factor can be very wide in a Power Meter. At higher differences between C and N , the impact on determination of the value of C is less significant but must still be considered. Using a Power Meter to make the measurements without a pre-detector filter would be impossible because it would be impossible to characterize N_0 across the entire measurement Bandwidth B . With a pre-detector filter one need only calibrate the N_0 value within the filter bandwidth.

A.3 Consideration of the Spectrum Analyzer approach

Spectrum analyzers are designed to make very accurate relative power spectral-density measurements. They do not exhibit great degrees of absolute power measurement accuracy, and must be periodically calibrated to ensure the best possible accuracy in such measurements. However, spectrum analyzers have other inherent advantages. One is that they can make very rapid measurements. While Power Meters have some specific settling time, which is based on the absolute RF energy levels arriving at their detectors, spectrum analyzers are designed for rapid swept-frequency measurements.

Spectrum analyzers can be considered to perform the type of measurements represented by half of Figure 8 A above in one box. They utilize an oscillator, operating either in fixed-tuned or frequency sweeping modes, to convert frequencies entering the input to the spectrum analyzer to a fixed IF frequency which then passes through a bank of selectable filter bandwidths to a fast power detector. In the type of measurements anticipated here, spectrum analyzers contain a range of available IF filters which can closely match the characteristics of the downlink signal. A very narrow filter has the effect of significantly reducing the Noise power component N contained in the $(C+N)$ measurement. The reason for that is that a narrow filter imposes a narrow bandwidth, B , on the $(C+N)$ measurements and from Equation 3 above the result is less total N energy in the $(C+N)$ measurement.

Spectrum analyzers make these measurements by collecting the individual detector power spectral-density measurements on a periodic basis. Typical Agilent spectrum analyzers, for example, record 700 or 1000 individual power spectral density measurements on each sweep. (An HP 8566 spectrum analyzer, for example, records 1001 measurement points in each sweep which can be as fast as 20 ms per sweep. The measurements are spaced evenly as a fraction of

the spectrum analyzer frequency span in scanning measurements, or even as a fraction of the total time in a time-based sweep.) These individual measurements are available on the screen and for output to a recording device. Additionally, spectrum analyzers utilize measurement “markers” which can quickly report the power spectral-density measurements of specific frequencies or time offsets.

On a spectrum analyzer one observes (C+N) spectral-density directly. Each one of the spectrum analyzer measurements described above actually represents the total (C+N) power contained within the spectrum analyzer Resolution Filter bandwidth at the frequency of the measurement. One could utilize two separate spectrum analyzers to independently measure the downlink signal from both the satellite of interest and the adjacent satellite. Such an approach would essentially simply replace the filters and Power Meters in Figure 8 A with separate spectrum analyzers. However, the approach in Figure 8 B goes one step further in simplifying the measurement configuration while actually improving measurement accuracy.

The approach illustrated in Figure 8 B takes advantage of the fact that spectrum analyzers have improved relative measurement accuracy over absolute measurement accuracy. It also takes advantage of spectrum analyzers’ inherent ability to make multiple quasi-simultaneous measurements with no degradation in measurement accuracy. Essentially, the two downlink signals in Figure 8 B arrive at the respective terminal downlink IF outputs at essentially the same frequency. (The only factor causing any potential differences in downlink IF frequency from each terminal would be differences in the terminal Block Downconverter Local Oscillator frequencies, different satellite transponder Local Oscillators, or different Doppler effects on the two satellites. The probability of differences being significant enough to differentiate the downlink signals in a spectrum analyzer measurement is very low.) The two downconverters in Figure 8 B are then tuned to slightly different frequencies. Such tuning can be accomplished by using slightly different 10 MHz reference input frequencies for the downconverters, as it will permit differences smaller than the typical tuning step size of synthesized downconverters. This step will permit the two signals to arrive at the spectrum analyzer at frequencies sufficiently different that the (C+N) measurement from one is not affected by the (C+N) energy from the other.

To permit both signals to be measured by a single spectrum analyzer, a power combiner is utilized ahead of the spectrum analyzer input. Under such conditions, the Noise energy from the two terminals combines to potentially degrade (C+N) measurement accuracy. If the N_0 energy from both terminals were exactly the same, the observed N energy would increase by 3 dB from that produced by a single terminal. If one were to be significantly less than the other, as in the (C+N) measurement discussion above, one would be dominant. However, if the sum of the two Noise power signals is still 10 dB below each of the (C+N) measurements, the impact on determining C would be low.

Figure 9 illustrates how a spectrum analyzer display of the two downlink signals, measured simultaneously, might appear. In this case, the desired signal is the lower frequency signal. It is measured on the spectrum analyzer display with a Marker on the signal peak—in this example at a frequency of 69.900 MHz and amplitude of 8.20 dBm. The second signal, from the adjacent satellite, is shown with a Delta Marker, in this case showing a Delta frequency of 200 kHz and Delta amplitude of -3.60 dBm.

To make accurate measurements under such conditions, it is important that the (C+N) energy of each carrier is clearly distinct, as illustrated in Figure 9, and do not overlap in frequency. There will already be some measurement contribution from the Noise energy of both downlinks, as described above, but if the (C+N) energy also overlaps in frequency it will be very difficult to subtract the adjacent C power to maximize measurement accuracy. Additionally, as described in previous sections the peak (C+N) power measurement should be at least 10 dB above the Noise Floor (approximately -4 dBm in this example) to ensure the effect of the Noise contribution to (C+N) is low. In the example, the Marker Delta shows approximately 8.4 dB in excess of the noise floor so the measurement setup should actually be changed to ensure that both (C+N) signal peaks are 10 dB or more above the Noise Floor to provide improved accuracy.

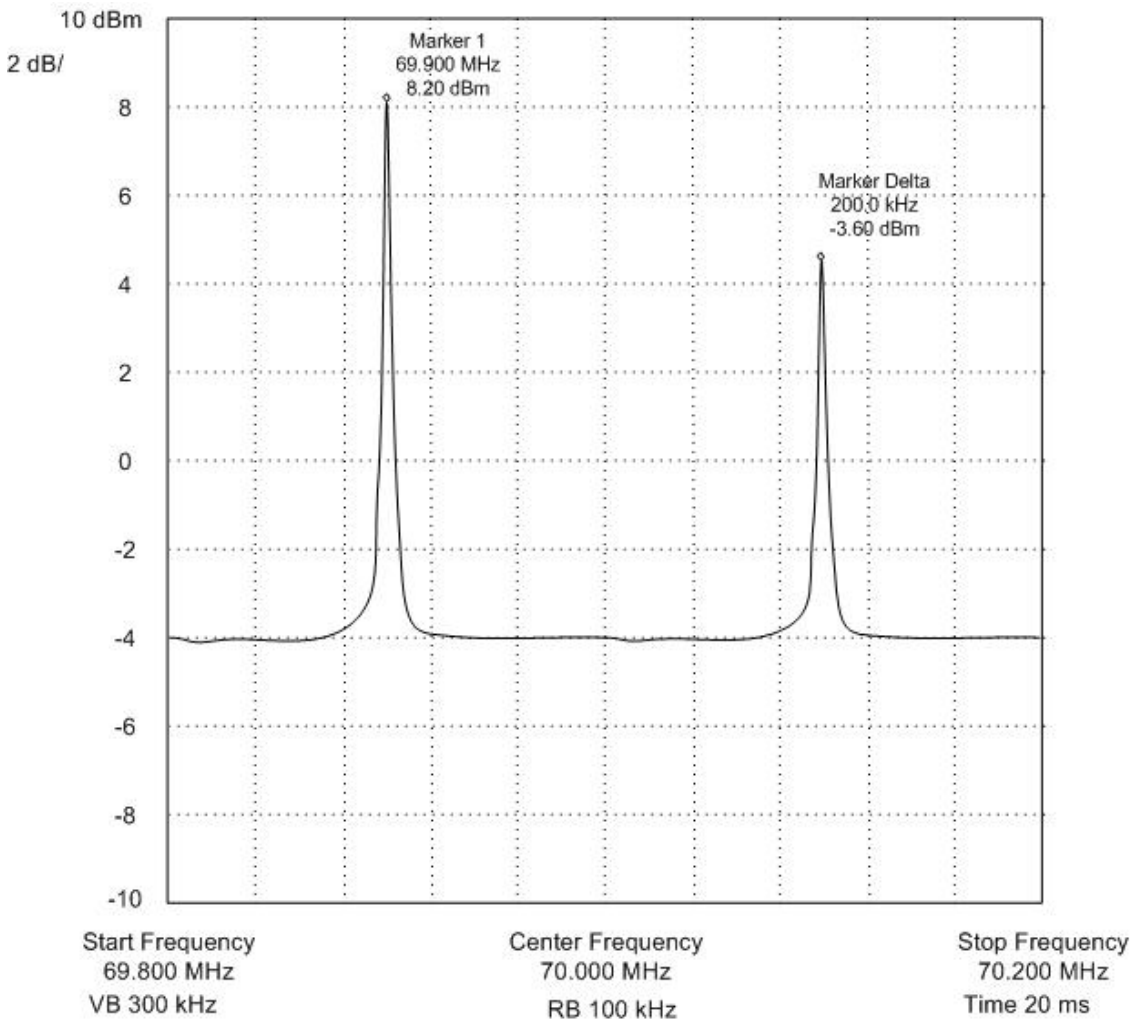


Figure 9 Simultaneous dual CW Signal Measurement Spectrum Analyzer Display

Utilizing this technique, simultaneous measurements of the (C+N) energy contained in the downlink from the desired satellite as well as (C+N) energy contained in the downlink from the adjacent satellite can be made. The results of the measurements could be recorded by a Data Collection Computer in two ways. The first approach would be to simply record the complete trace data from each sweep of the spectrum analyzer. This would result in the accumulation of a large amount of recorded data for relatively little measurement information, since the measured (C+N) values of each carrier are all that are really required. However, it would provide a record of the spectrum actually being observed that might preclude any doubts about (C+N) above the Noise Floor and interference issues. Additionally, the time required for the spectrum analyzer to report the complete trace data to the Data Collection Computer is often significantly longer than the actual sweep time. In such cases, the test and measurement system would be limited by the spectrum analyzer / Data Collection Computer communication time rather than the actual spectrum analyzer sweep time.

The alternative approach would be to record simply the Marker and Delta Marker information from the spectrum analyzer. This is a much smaller data set for each measurement point and is thus much faster to communicate from the spectrum analyzer to the Data Collection Computer. It would thus be limited simply by the spectrum analyzer sweep time and not by the spectrum analyzer / Data Collection Computer communication time. The only drawbacks to this approach are that there would be no absolute record that each measurement resulted in (C+N) values sufficiently above the Noise Floor to ensure maximum accuracy, and there would be no record that interference did not affect the measurements. Still, an operator conducting the

measurements could ensure through observations prior to the commencement of data collection that these two desired conditions are met. The spectrum analyzer will continue to display the full sweep data visually during the test process, so ongoing visual observation could confirm that the desired measurement conditions were met.

An additional measurement simplicity is also offered by this measurement approach. If the spectrum analyzer is set to "Max Hold" during the measurement period, it will record the highest signal level for each "bin" in the continuous spectrum analyzer sweeps during the measurements. Thus, the highest values observed by the spectrum analyzer for the (C+N) of the satellite of interest downlink signal as well as the (C+N) of the adjacent satellite downlink signal will be recorded. This would give an absolute maximum tracking error signal indication because it would record the highest level observed in the downlink of the adjacent satellite. For measurements in which it is desired to ensure that antenna pointing and tracking are within some preset limit, this would provide absolute assurance of such conditions without even the need for a Data Collection Computer. For conditions of considerable variation in either or both of the downlink signals, this measurement may be of little value. However, for the normal VMES operating case where antenna tracking is expected to be extremely accurate, it can provide a quick "Go / No-Go" confirmation of satisfying tracking performance objectives.

A.4 How to make accurate CW signal measurements using a Spectrum Analyzer

Since a spectrum analyzer will be used to make the downlink (C+N) measurements, it is imperative that the tests be conducted by an operator who knows how to make accurate spectrum analyzer measurements under these conditions. Specifically:

- The downlink signals will consist of essentially CW tones.
- Peak (C+N) values will be 10 dB or more above the Noise Floor of the measurement setup.
- Peak (C+N) values will be far enough in frequency that the presence of C energy from one carrier will have negligible impact on the C energy measured from the adjacent carrier.
- The spectrum analyzer will perform swept frequency measurements at sweep speeds consistent with the spectrum analyzer maximum measurement accuracy.

Spectrum analyzers make power spectral-density measurements by tuning an oscillator such that a precise IF filter can be used to select the bandwidth of power spectral-density that is measured by the spectrum analyzer detector. In conducting this operation, there are several significant factors which affect measurement accuracy.

The IF filter used by the spectrum analyzer to select the increment of RF spectrum reaching the detector is called the "Resolution Bandwidth" filter. It must be very precisely measured and controlled to ensure spectrum analyzer measurement accuracy and repeatability. Spectrum analyzers typically have internal calibration routines which use a CW tone input to the spectrum analyzer which then adjusts the shaping and amplitude of the raw measurement data based upon the designed Resolution Bandwidth characteristics. After the spectrum analyzer has run its internal calibration routine, it should be able to make very accurate power spectral-density measurements that are repeatable from one Resolution Bandwidth to another. Agilent spectrum analyzers typically contain a range of Resolution Bandwidth filters in 1,3, 10 bandwidth increments. For example, a particular Agilent spectrum analyzer might have Resolution Bandwidths of 10 Hz, 30 Hz, 100 Hz, 300 Hz, 1 kHz, 3 kHz, 10 kHz, 30 kHz, 100 kHz, 300 kHz, and 1 MHz available. The bandwidth selected for CW measurements is critical because the highest sensitivity of (C+N) measurements above the Noise Floor occur when the narrowest possible Resolution Bandwidth filter is employed. Even though CW signals are essentially 0 Hz wide, the spectrum analyzer may not be able to use its smallest Resolution Bandwidth filter to make the desired measurement due to drift in the frequencies of conversion oscillators anywhere

in the system of interest, or in the spectrum analyzer itself. Additionally, a very narrow Resolution Bandwidth filter takes longer for the detector to make a measurement than a wider filter, unless Digital Signal Processing (using Fast-Fourier Transforms, or FFTs) are employed. For those two reasons it is usually prudent to select a Resolution Bandwidth filter as small as possible while still providing reasonable sweep times and tolerance for combined frequency inaccuracies.

After the spectrum analyzer detector, there is another filter which provides “smoothing” of the measurement result. In Agilent spectrum analyzers it is called the “Video Bandwidth” filter. A narrow Video Bandwidth filter tends to smooth out the random nature of noise effects on signal measurements. It has the effect of essentially “averaging” the detector output for a small time period following each measurement across the sweep. Wide Video Bandwidth filters permit the spectrum analyzer to observe very rapid signal fluctuations, such as modulation, which can affect the measurement data. Smaller Video Bandwidth filters require more time for the spectrum analyzer to pause and smooth the results after each detector measurement so they will slow the spectrum analyzer sweep speed. When measuring CW tones in the presence of a Noise Floor significantly below the (C+N) energy levels, it is best to select a Video Bandwidth that permits faster sweep times while still providing a noticeably “smooth” Noise Floor. If the Video Bandwidth filter is selected at too broad a bandwidth, the apparent Noise Floor might have fluctuations of several dB and thus impose random variations in the resultant C measurements when N is subtracted from (C+N).

Another factor in Agilent and other spectrum analyzers is Video Averaging. That technique is quite similar in results to narrowing the Video Bandwidth, but it is achieved by averaging multiple sets of spectrum analyzer sweep data, point by point. For example, if Video Averaging is set to 4, 4 separate spectrum sweeps are performed and the data at each point in the spectrum analyzer display is averaged for the measurements at that exact frequency point in each of the 4 sweeps. As additional sweeps are performed, the spectrum analyzer continues to display a “running average” calculated by averaging the prescribed number of previous sweeps. Video averaging can effectively flatten the Noise Floor and permit reliable measurements of Noise energy in a way which is faster than using a narrow Video Bandwidth, and that is the main reason it was developed. For the purposes of our measurements here, we should be able to make reasonably fast measurements using a suitably small Video Bandwidth filter rather than resorting to Video Averaging which could degrade the absolute (C+N) measurements.

Finally, spectrum analyzers typically conservatively calculate the sweep time requirement to make a particular measurement based on the operator-selected values of Resolution Bandwidth and Video Bandwidth. It does that by calculating the time required for the spectrum analyzer detector to accurately respond to the input signal power spectral-density given the established filter settings. In most spectrum analyzers this internal calculation can be “over-ridden” by direct operator input of Sweep Time. There are some cases in which this might make sense, such as when observing the signal modulation characteristics rather than making accurate power spectral-density measurements. However, CW measurements as described here are not such a case. For the maximum CW measurement accuracy, the spectrum analyzer’s internally-calculated Sweep Time should be used, and not forced by operator intervention. Most spectrum analyzers will report a special “Uncalibrated” message when a Sweep Time below that calculated by the spectrum analyzer is selected. It is important for best measurement accuracy here that such messages are not generated during operation.

Figure 10 illustrates a spectrum analyzer display for a representative CW signal measurement. In this example, the signal of interest is centered at 70 MHz and Marker 1 displays its measurement frequency of 70.000 MHz and amplitude of 8.01 dBm. It should be recalled for the earlier discussion that the value of 8.01 dBm actually represents the (C+N) valued observed by the spectrum analyzer, and not the value of C itself.

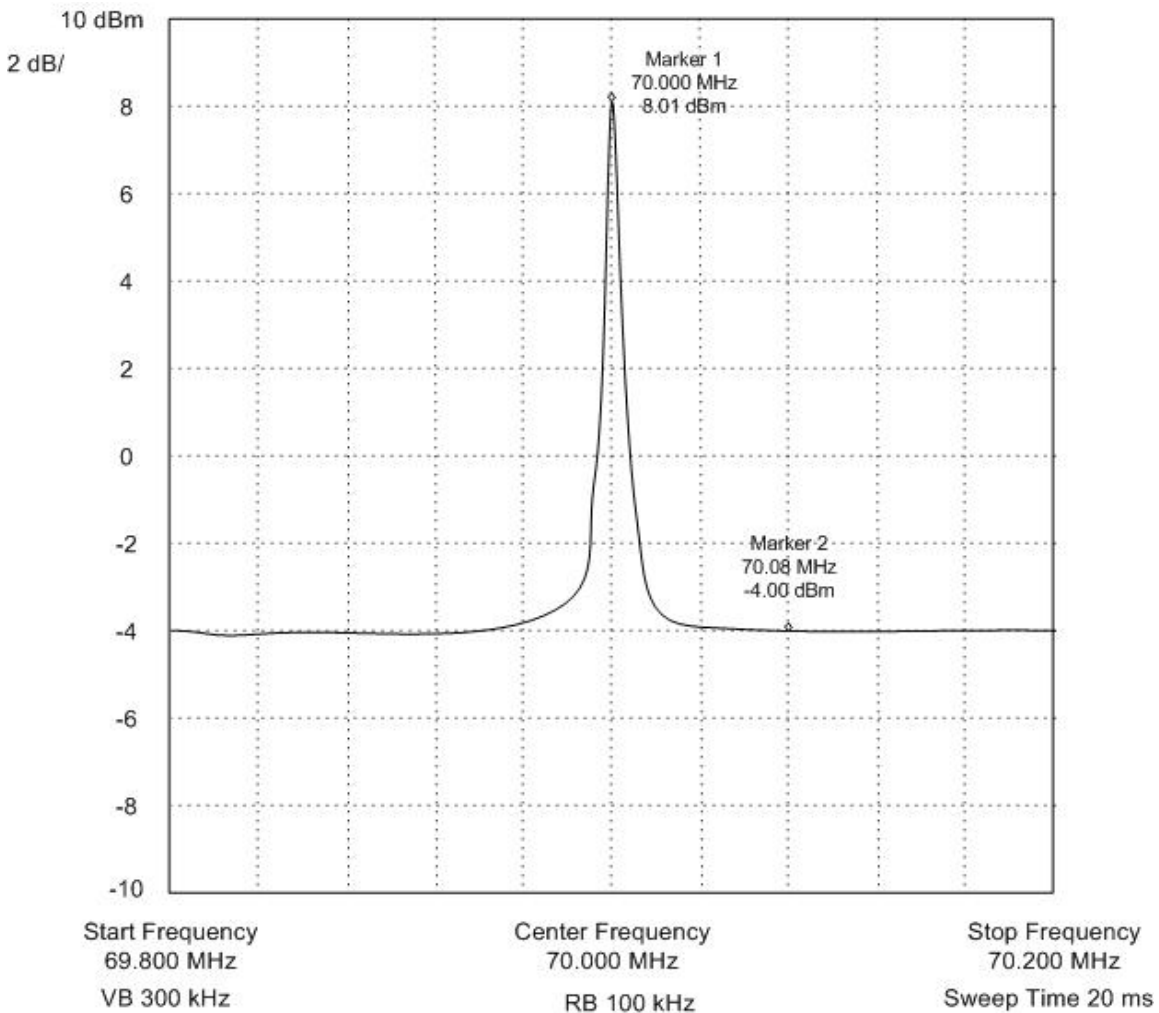


Figure 10 Single CW Signal Measurement Spectrum Analyzer Display

To be able to make an accurate measurement of C, it is necessary to have a reference of the N value contained in the (C+N) value so it can be mathematically subtracted. In the example of Figure 10 this Noise measurement is made with Marker 2 and shown at frequency 70.08 MHz and amplitude -4.00 dBm.

One of the considerations when making such measurements is where to make the Noise measurement. The value of N needed is that exactly at the frequency of the (C+N) measurement. Since we can't perform the measurement at that exact frequency, we select one as close as possible which will represent the apparent Noise value. In the case of Figure 10 the Noise Floor is relatively flat so an N measurement could be made almost anywhere in the spectrum analyzer trace outside the (C+N) represented by the carrier. In cases where the Noise Floor has an apparent slope, the most accurate N measurements are actually made by assuming the Noise Floor to be smooth in the dB domain, and making measurements on each side of the carrier of interest and then averaging the two.

Having collected the measurement data represented by Figure 10 one could then calculate the amplitude of the Carrier Power, C. To do so, we utilize Equation 3 above, which is:

$$C = (C+N) - N_0 \times B \quad (\text{Equation 3})$$

The essential steps in calculating accurate values of C from the spectrum analyzer observations are to:

- 1) Convert (C+N) values to linear power units
- 2) Calculate N by multiplying N_0 over the Resolution Bandwidth B, or use the N value from the spectrum analyzer directly
- 3) Convert N into linear power units
- 4) Subtract N from (C+N) to determine C, in linear power units
- 5) Convert C from linear power units to dBm or dBW

These steps will be performed below using the measurements in Figure 10 as an example.

Marker 1 represents (C+N), but in log units of dBm. Converting (C+N) from log (dBm) units to linear power units we determine that:

$$8.01 \text{ dBm} = 6.324 \text{ milliwatts} = (C+N)$$

Spectrum analyzers can measure the Noise Floor by setting a marker, as shown in Figure 7, or the "Marker Noise" function can be selected on the spectrum analyzer in which case it will calculate N_0 directly from the measurement. In the example of Figure 10 we conclude that the marker 2 measurement is the total N power contained in one Resolution Bandwidth of the spectrum analyzer. Thus, it can be used directly and we must simply convert it to linear power units. We calculate that:

$$-4.00 \text{ dBm} = 0.3981 \text{ milliwatts} = N$$

Now, we calculate C by subtracting N from (C+N) as described in Equation 3.

$$C = 6.324 \text{ milliwatts} - 0.3981 \text{ milliwatts} = 5.926 \text{ milliwatts}$$

Converting to dBm:

$$C = 5.926 \text{ milliwatts} = 7.728 \text{ dBm}$$

It would be this value, C, that would be utilized in the antenna pointing measurements to compare the changes in downlink power that would be expected for a given antenna pointing error.

A.6 Factors which might affect signal measurement accuracy

Several different factors should be considered which might affect the accuracy of the measurements obtained. Some of these are a function of the test equipment setup and can be controlled while establishing test conditions and others are environmental which should be monitored during the period of the test. The description below will consider these potential factors and propose methods to constrain their affects.

- 1) Propagation impacts. Obviously, VMES operations are susceptible to all the same propagation effects as all FSS Ku-Band signals. Additionally, the VMES environment produces a large number of signal blockage conditions when antenna look angles are obstructed by buildings, utility poles, vegetation, and other objects. Careful VMES testing should only be performed under reasonably "clear sky" conditions to ensure rain fades or other similar propagation degradations are not encountered. Signal blockage cannot reasonably be avoided in VMES operation, so about the only mitigation for these effects is to attempt to utilize conditions which would block the signal path to the target satellite and adjacent satellite used for the test in the same way. This can likely be best confirmed if suspicious results appear in the measurement data. For example, in most VMES blockage conditions a vehicle would pass from "clear" to "blocked" conditions in a way that would result in downlink levels on the target and adjacent satellite about equally. If the test data suggests this not to be the case, consideration should be given to the exact blockage conditions which might produce such results.

- 2) Vehicle dynamics. VMES terminals are designed to satisfy operational requirements under a specific set of dynamic conditions. While these conditions are not specified by the FCC, they are often specified in customer requirements documents or specifications. Such levels of vehicle dynamics are actually one of the differences in the design of different VMES systems and can serve as market differentiators. VMES testing can be performed using carefully instrumented vehicle dynamics or it can be performed using available test vehicles, road surfaces, or “test track conditions”. For repeatability, VMES testing should be performed in some controlled way such that the vehicle dynamics are either carefully measured or repeatable. Excessively high dynamics should be expected to have the result of degrading VMES antenna pointing test data.
- 3) Satellite transponder linearity. Measuring antenna offset angles in the approach described here presumes that the transponders on both the desired and adjacent satellites are operating in a linear mode. Thus, efforts should be made to ensure that signal levels on both satellites are significantly below saturation and well within the linear region. Additionally, efforts should be made to confirm that during the period of the test there are no unusual satellite transponder operations like gain state changes, transponder signal level fluctuations that might be caused by high-level TDMA signals, etc. Linear operation can be confirmed by observing transponder power prior to beginning the test. Transponder gain stability can be confirmed by a period of signal monitoring prior to beginning the test.
- 4) Test system linearity and stability. Just as transponder linearity can affect the test data, the test setup itself can also affect measurement accuracy. All amplifiers and frequency converters used in the test setup must be confirmed to be operating in their linear range and without any variations in gain or frequency. For maximum accuracy the test setup should be in operation for a considerable period of time before actually beginning the tests to ensure that all associated equipment has stabilized in temperature and frequency prior to making any measurements. Frequency sources in every converter should be confirmed to be phase-locked and stable. Gain of all amplifiers and frequency converters in the test setup should be confirmed stable and free from any significant fluctuation in temperature, primary power, or other factors. It should be confirmed by monitoring that during the entire test period the frequency of the downlink signals both on the desired and adjacent satellites have not drifted enough to degrade spectrum analyzer measurements. (This would likely be most significant if the spectrum analyzer were to be used in a fixed Marker mode. In such cases the spectrum analyzer marker stays on exactly the frequency set at the start of measurements. Frequency drift farther than a single Resolution Bandwidth filter bin in the spectrum analyzer would then cause the Marker to produce faulty measurements.) Obviously, no changes in any element of the test setup should be made while measurements are actually being collected.
- 5) Test signal levels sufficiently above the Noise Floor. Since the test measurements require accurate determination of C power above the satellite transponder and test equipment Noise Floor, steps should be taken to ensure RF and IF power levels are sufficient for measurement. This should be confirmed through multiple verifications prior to conducting the test. Specifically:
 - a. To ensure the spectrum analyzer noise floor is not affecting measurements, after the test setup is configured, including the specific spectrum analyzer frequency, amplitude, sweep, and filter settings, the input from the spectrum analyzer can be temporarily disconnected. The Noise Floor on the spectrum analyzer in the absence of any input signal should be confirmed to be more than 10 dB below the Noise Floor for the test signals themselves.
 - b. The G/T of the downlink test terminals can degrade measurement accuracy if it is not high enough to make reliable Noise measurements. If the two

downlink test terminals have differing G/T performance, the terminal with the better G/T should be utilized on the adjacent satellite rather than the desired satellite because the adjacent satellite will operate at lower absolute signal levels. Each downlink signal should be temporarily removed from the power combiner ahead of the spectrum analyzer and expected changes in measured Noise Floor confirmed. (If the G/T of the downlink test terminals, removing either input should cause a 3 dB decrease in measured Noise Floor. If one is significantly better than the other, removal of the signal from the terminal having the higher G/T should have relatively little impact on the measured Noise Floor.)

- 6) Inaccuracies in antenna patterns. The test technique utilized here presumes that the VMES transmit antenna radiation patterns are performing as previously measured. There are several factors which might cause the measurements to diverge from the previous antenna pattern measurement conditions. These include such factors as:
 - a. Antenna patterns are measured in the azimuth and elevation planes. The desired and adjacent satellites will rarely be in exactly only the azimuth plane. Circular VMES antennas will have radiation patterns that are similar in all planes. Non-circular VMES antennas will have radiation patterns that could exhibit differences in planes that further pattern estimation may well be required.
 - b. Radomes mounted on the VMES antennas will have some impact on the antenna radiation patterns. As a first approximation, the tests describe here will disregard the radome effects. Depending upon the results obtained, it may be prudent to run a series tests with the radome installed and another with the radome removed to quantify the potential impact on radiation patterns.

Appendix B: List of acronyms

Acronyms utilized in this document are included in this appendix.

<u>Acronym</u>	<u>Definition</u>
ASI	Adjacent Satellite Interference - RF Energy radiated towards satellites adjacent to those intended for communication
ATE	Authorized Test Entity - An entity which performs verification tests in accordance with GVF authorization
BPE	Beam Pointing Error - error in earth terminal antenna pointing direction from desired direction
COTM	Communications On The Move - VSAT terminal capable of operating while mobile as well as in a fixed location.
CW	Continuous Wave
EIRP	Effective Isotropic Radiated Power
EIRPSD	Effective Isotropic Radiated Power Spectral Density - EIRP within a specific measurement bandwidth
ETSI	European Telecommunications Standards Institute
FCC	United States Federal Communications Commission
FFT	Fast Fourier Transform
GSO	Geo-Stationary Orbit
IF	Intermediate Frequency
RF	Radio Frequency - in this context, satellite uplink or downlink frequencies
TDMA	Time Division Multiple Access - an RF transmission scheme in which multiple transmissions are separated in time on a given RF frequency.
UPC	Uplink Power Control - an automatic method of controlling uplink power in an earth terminal transmitter to compensate for changes in link losses, antenna radiation patterns, or other effects.
USAT	Ultra Small Aperture Terminal
VMES	Vehicle-Mounted Earth Station
VSAT	Very Small Aperture Terminal